Microlensing in the Era of All Sky Surveys

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Measurable quantities

- Einstein crossing time t_E duration $t_E \propto \frac{\sqrt{M \pi rel}}{\mu rel}$
- Microlensing parallax $(\mathbf{\pi}_{E})$:
 - Over the course of the year observing microlensing event from different angles can lead to bumps and asymmetries in the lightcurve
- Identify black holes in $\pmb{\pi}_E \, vs \; t_E$





 $\sqrt{\pi}rel$

 $\mathbf{\pi}_{\mathrm{E}} \propto$

Need large photometric survey to find black holes through microlensing



Rubin Observatory Community Driven Survey Optimization Process



Quantifying Impact of Rubin Survey Strategy on Microlensing

- Focusing on high stellar density areas (e.g. Galactic plane and bulge) have the most significant impact on microlensing detection and characterization
 - + Including them improves BH event detection by ${\sim}50\%$
- Rubin will measure BH population vs place in galaxy
 - Taking observations throughout the plane optimizes ability to discover events throughout the plane with a 10-20% drop in characterization efficiency
- Rolling in the bulge and plane should be done carefully to allow for microlensing parameter characterization
 - Roman or less severe rolling (optimize to fill in gaps)





Testing methods with ZTF data First three years of ZTF: 60 Events (19 outside of galactic plane)



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2022.

Simulated vs OGLE vs ZTF $t_{\rm E}$ distributions in good agreement



No likely black hole candidates in first 3 years of data





Medford & Abrams et al. 2023

802 Ongoing Candidates with t_E > 150 days

- No completed black hole candidates
- If same rate of real microlensing events as complete events at this stage → a few candidate black holes
- In 5 years of data found 124 events with 12 ≥ 150 days (Zhai et al. 2023)





Investigating Nature of Out of Plane Objects

- 5 events have Gaia parallaxes with significance > 3σ
 - 0.7-4 kpc away
- 3 events have SDSS redshifts z ~ 1, 2 of which are classified as quasars (AAVSO and Million Quasar Catalogs), but also have Gaia parallaxes
- 2 events have radio signals < 1" away (NVSS)
- In Progress: Follow-up data to verify Galactic origin of event







Predicting and Modeling Galactic Black Hole and Binary Populations with PopSyCLE Simulation

PopSyCLE

- Population Synthesis code
- Useful beyond microlensing!

Galaxia – Sharma et al. 2011; SPISEA – Hosek et al. 2020; PopSyCLE – Lam et al. 2020

https://github.com/jluastro/PopSyCLE



We want to know how often black holes are in binaries and understand when they make a measurable binary signal or when they masquerade as single lenses



More than half of stellar and compact object microlensing events have at least one binary $...\sim52\%$ appear as singles



Abrams et al., in prep

Brings t_E distribution into agreement with data 10^4

- Mean $t_{\rm E}$ averaged across 18 fields:
 - OGLE: 25.6±3.3 days
 - Singles only: 18.7±1.2 days
 - Singles + single peaked multiples: 25.6±1.4 days

- S Run Model with only singles
- M Run Model with multiples and singles

Mróz et al. 2017 Mróz et al. 2019



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Binary Parameter Bias with Microlensing





Abrams et al., in prep

Summary

- Inclusion of high stellar density areas in Rubin such as Galactic bulge and plane increase microlensing discovery and characterizations by ${\sim}50\%$ for black holes
- We discovered 60 high-quality microlensing events in the 3 yrs of ZTF data including 19 outside the Galactic plane
- We find that > 50% of our simulated events with PopSyCLE include a binary lens or source system creating both false positives and false negatives when identifying BHs
- Critical for microlensing surveys to include binarity in predictions and modeling



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ZTF Paper: Medford & Abrams et al. 2023