The epic of EROS Expérience de Recherche d'Objets Sombres

30 years of research on microlensing and variable stars

this talk is dedicated to those of EROS who have passed away

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M. Moniez (IJCLab-CNRS)

Last minute: At the start of EROS, I heard almost the same remark as J-P Beaulieu (see his talk), but from another distinguished professor: « *You will never discover microlensing events...* »

Instruments, detectors

EROS-1: 1990-1994

- 290 photographic plates taken at the 1m Schmidt (ESO), digitized with the MAMA (1Gpixels/plate). 5.10⁶ stars.
- (first) 16 buttable CCD-camera (4M-pixels!) on a 40cm telescope. 19,000 images of the LMC bar. 2.5.10⁵ stars

EROS-2: 1996-2003 2 cameras with 32 Mpixels each, on a 1m telescope



Main targets

monitored by EROS2 in 1996-2003 (~ 40 participants)

Magellanic Clouds => probe hidden matter in halo (t ~ 5.10⁻⁷)

Galactic center => probe ordinary stars as lenses in disk/bulge (t ~ 2.10⁻⁶)

Spiral arms

=> probe ordinary stars in disk, bar + hidden matter in thick disc (t ~ 5.10⁻⁷)

50 Terabytes of data

850,000 images processed ~ 77x10⁶ stars measured 300 to 500 times

The first discoveries: 1993 => signal ≤ 2 events





1998: The first caustic crossing monitored around the world

EROS contribution in Afonso et al. A&A 337 (1998) L17





1998: The first caustic crossing monitored around the world





My personal memories of this episode

Big Brother 1996 1997 1998 1999 2000 2001 2002 2003 construction

Mercredi 17 Juin 1998

- I had **just arrived at La Silla** with a team of mechanical engineers (C. Bourgeois & R. Cizeron) for a major intervention on the telescope
- Weather, which had been rainy the previous day, suddenly improved just before SMC became visible (1h30 am)
- I forced all the protection limits to point the telescope as soon as possible
 -> SMC elevation was 26°52'
- For 5 hours, extraction of **photometry in real time** enabled to properly **adjust exposure times**.

In those days, we still used landline telephones to communicate with South Africa and Australia, so as to be kept precisely informed of the progress of the event and observations. 20h18: Shift J.H. Ch.C + Marc et al. arrives cet apres midi. 20h19: 20h10: Tres mauvals temps toute la journee: plages et vent. 20h20: Sur ta rin, incluing jouge 0 02%. Pas de flats. 20h21: Nous sommes prets pour la "nuit caustique", avec le Marly, 20h22: le Danois, et une T.O. pour un spectre NTT. 20h22: Le ciel aura peut etre pitie de nous: quelques etoiles sont 20h23: visibles dans un ciel encore perturbe, Mais l'humidite 20h23: se maintient autour de 80%

 of photometry in real y adjust exposure times.
 01048: Lancement de smc207 a 18h24 ST. 01h39: De Noull 1 De N



Many complications to Point-source-Point-lens rectilinear events



EROS combined results in dark matter

(1) Galactic halo: all data + combine EROS+MACHO toward LMC









EROS combined results in dark matter (1) Galactic halo: all data + combine EROS+MACHO toward LMC 1.0 0.8 Masses in the Stellar Graveyard a major $f = f / 4.7 \times 10^{-7}$ Galacti 0.6 nsing Φ Ο ō Now exclude component 5 Ε 6 0.2 0.0 0.001 0.100 10^{-5} 10 1000 10^{-7} M/M_{\odot}

Combining light-curves to search for t_E>100 days events





detection efficiency for events > 100 days greatly improved by combining light curves



Combining light-curves to search for t_E>100 days events





detection efficiency for events > 100 days greatly improved by combining light curves



Note: C. Stubbs offered to use superMACHO data too, but the superpowers of a Harvard professor weren't enough to overcome bureaucracy

(2) Galactic plane: Galactic center



(2) Galactic plane: Galactic center

The optical depth t - probability for a source to be behind an Einstein disk at a given time (Amplification > 1.34) [Adapted from Specht et al. (2020)] $MaB\mu IS2$ all sources Mapping of t in MOLTHAN 6 **OGLE-IV** all sources galactic latitude MOA-II RCG sources 5 Incon fit (all) MOA-II all sources Fit (RCG) $\tau \begin{pmatrix} \times 10^{-6} \\ \infty \end{pmatrix}$ 2 0 b (deg)

(3) Galactic plane : Galactic spiral arms (only EROS before 2020)



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Other science

• Systematic search for SNIa (1999 and 2000) measurement of explosion rate @z = 0.13 within the current uncertainties (2±0.6.10⁻⁵/yr/cMpc³), based on 16 confirmed SNIa: *Blanc et al. A&A 423, 881 (2004)*

• Periodic variable stars

d-Sct, cepheids, RR-Lyrae, eclipsing binaries... within LMC: *Dae-Won et al. A&A, 566 (2014)*

- **Rarities (non-periodic):** R-coronae borealis, caustic crossings (microlensing again)
- **SN1987a echoes:** found as a source of fake microlensing when superimposed with cataloged stars
- **Potential** for asteroids, EROS LAC deep-field...



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Serendipitious discovery

Residual microlensing background



SN1987A echoes

Stars bumped

SN1987a

- Diffusion of the light emitted by the SN on interstellar gas
- If superimposed on a monitored star

1st flash received directly

the 23rd february 1987

- -> brightness seems to vary
- Unexpected microlensing background...

Reliable results come from a deep understanding of the detector => Not frequently shown because fiducial is not good com.



EROS: 10 years of data 30 years of analysis



- Dark matter results exceed expectations: exclusion from 10⁻⁷ to 1000xM_{sur} Ο
- Galactic Center, Galactic Spiral Arms (EROS exclusivity until 2020) Ο
- Extras: Caustic-crossings / microlensing exotics Ο
- Other results + searches : all types of variability
 - In 1998, EROS was the main discoverer for SN's @z ~ 0.1 0
- **Rarities**
 - R-coronae Borealis, DY Per stars 0
 - SN1987a echoes 0



EROS legacy, a brilliant future: EROS is finalizing a database for public use

- Will be used by brokers (like Fink), to help categorization of the Rubin-LSST alerts. Ο
- For combined analysis with data from current and future surveys (OGLE, Gaia, ZTF, 0 Rubin-LSST, ROMAN...) for very long timescale variabilities

Few key numbers

- $O(10^8)$ stellar sources monitored for 10 years with 300-500 photometric measurements
- O(10⁵) stellar sources monitored for almost 20,000 times in 4 years (sampling 20 min.) •
- $O(10^4)$ cepheids and RR-Lyrae + variables of undetermined type (like LPV > years) •

Final final... in 50000 years... THE MESSENGER Archeological remains from EROS **2**



A time capsule whose contents are described in Dr. Richard West's article was then deposited by President Frei with the works being blessed by the Archbishop of Antofagasta, Monsignor Patricio Infante.

The Time Capsule

During the December 4, 1996 event at Paranal, President Frei placed a Time Capsule with selected materials in the wall of the UT1 building. This capsule, an aluminium cylinder of 15 cm diameter and 45 cm long, was filled with nitrogen gas and sealed hermetically. A commemorative plaque was fixed in front of the cavity. In addition to the various papers listed below, the capsule also contains a list of contents etched on a metal plate which will survive virtually indefinitely in this environment, while it cannot be excluded that the papers may deteriorate with time.

Contents of the Time Capsule:

- 1. Document signed by President Frei on the occasion;
- 2. Copy of the Acuerdo between ESO and Chile;
- One outstanding scientific paper from each of ESO's member countries (a list of the titles is available at the ESO Web-site)
- Copy of December 4 issues of the Chilean newspapers El Mercurio and La Epoca;
- 5. Copies of the written version of the speeches delivered on the occasion;
- 6. Viewgraph and description of VLT, as presented at the exhibition;
- 7. List of VLT Team members;
- 8. Photos of Paranal before and after construction;
- 9. Photo of the VLT model at exhibition.

R.M. West

Final final... in 50000 years... THE MESSENGER Archeological remains from EROS ?



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hoping to see you for the 50th anniversary (before the 50000th). I'm saving a few other anecdotes for this occasion

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